

# Practical Horizon Plane and Communication Duration for Low Earth Orbiting (LEO) Satellite Ground Stations

SHKELZEN CAKAJ

Post and Telecommunication of Kosovo (PTK),  
Dardania, nn., 10000 Prishtina,  
KOSOVO

E-mail: [Shkelzen.cakaj@ptkonline.com](mailto:Shkelzen.cakaj@ptkonline.com); [scakaj@yahoo.com](mailto:scakaj@yahoo.com)

**Abstract:** - Communication via satellite begins when the satellite is positioned in the desired orbital position. Ground stations can communicate with LEO (Low Earth Orbiting) satellites only when the satellite is in their visibility region. The visibility region is in fact the horizon plane. Because of natural barriers or too high buildings in urban areas, practical horizon plane differs from the ideal one. The duration of the visibility and so the communication duration varies for each satellite pass at the ground station, specifically for LEO satellites which do move too fast over the Earth. This paper discusses the satellites motion detection, the difference in between ideal and practical horizon and further the variations of the communication duration between the ground station and LEO satellites. Main objective is determination of practical horizon plane and critical maximal elevation angle related to communication duration. For this paper, data recorded at the Vienna satellite ground station within the Canadian space observation project "MOST" (Micro variability and Oscillations of Stars) are applied. Vienna ground station system was set up at the Institute for Astronomy of the University of Vienna in cooperation with the Institute of Communications and Radio- Frequency Engineering of the University of Technology in Vienna.

**Key-Words:** - LEO, satellite, ground station, elevation, azimuth, horizon plane, communication duration.

## 1 Introduction

The typical satellite communication system comprises of a *ground segment*, *space segment* and *control segment*. The link which transmits radio waves from the ground station to the satellite is called *uplink*. The satellite in turn transmits to the ground station by the *downlink*. The function of the *ground segment* (one or more ground stations) is to receive or transmit the information to the satellite in the most reliable manner while retaining the desired signal quality. The *space segment* consists of one or more artificial satellites. In case of more satellites they are organized in a network called *constellation*. The *control segment* consists of all ground facilities for control and monitoring satellite. Ground stations are distinguished by their size which varies according to the volume of traffic to be carried and the type of traffic (voice, video or data). The largest ground stations are equipped with antennas of 30 m diameter (Standard A of the INTELSAT network). The smallest ground stations have typically 0.6 m antennas (direct television receiving stations). Some stations, both transmit and receive, and some of them are receive-only (RCVO) stations. The general organization of a ground station consists of antenna subsystem with

associated tracking system, transmitting and receiving equipment, monitoring system and normally power supply. Fig. 1 shows typical architecture of a ground station for both receiving and transmitting. The antenna is common to transmission and reception by reasons of cost and bulk. The separation of the transmission and reception is achieved by means of duplexer [1].

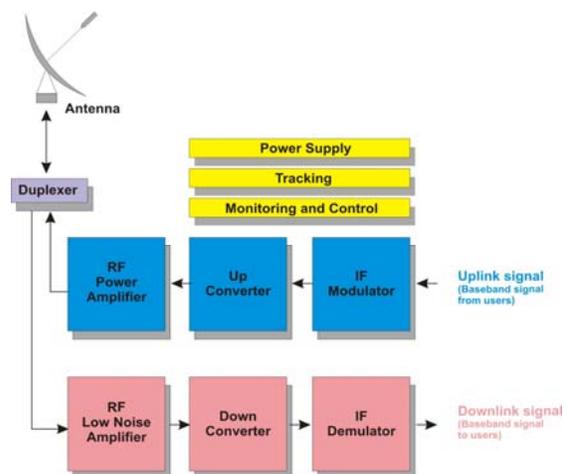


Fig. 1. The satellite ground station architecture

The *visibility region (horizon plane)* of the ground station is defined as a region on the Earth from where the satellite is seen from the ground station under the lowest elevation angle. The ideal horizon plane is considered under  $0^\circ$  angles of elevation. The communication between the satellite and a ground station is established only when the satellite is visible from the ground station. For low cost LEO satellite ground stations in urban environment it will be a big challenge to ensure communication down to the horizon [2], [3]. The communication at low elevation angles can be hindered through natural barriers or too high buildings in urban areas. Misspointing is another reason of loss in communication [4]. Thus, motion (appearance) detection of the satellite above natural barriers enables the practical horizon to be determined. Practical horizon differs from the ideal horizon. This difference impacts on communication duration in between the satellite and ground stations, specifically under low elevation angles.

Thus, in order to determine the practical horizon plane and its impact on communication duration between the satellite and ground station, the general concepts of artificial satellites, orbits, tracking of satellite and ground satellite station geometry are further given. Then, data recorded at Vienna satellite ground stations are applied to determine the practical horizon plane and maximal elevation angle under which communication between the ground station and satellite rapidly fall consequently causing loss in data transfer between the source and destination.

## 2 Artificial Satellites

An artificial satellite is manufactured object dedicated to continuously orbit the Earth, or other body in space. The original objectives of artificial satellites were to serve low-cost communications relays and to provide new opportunities on investigation and development of new radio techniques. Recently, especially with escalating cost of large satellites, attention is turned to smaller satellites so called *microsatellites*, which are taking also a new role, including science missions [5]. LEO satellites have very wide applications, from remote sensing of oceans, through analyses on Earth's climate changes, Earth's imagery with high resolution or astronomical purposes. These satellites provide opportunities for investigations for which alternative techniques are either difficult or impossible to apply. Ground stations have to be established in order to communicate with such

satellites. Scientific missions can be accomplished in principle by only one ground station. The reason behind building more ground stations is to increase the coverage and number of measurements per observed object or area, and practically increase data download capability [2], [6], [7].

An artificial satellite essentially consists of two main functional units: *payload* and *bus (platform)*. The primary function of the *payload* is to provide communication by repeater and antenna system. The *bus* provides all the necessary electrical and mechanical support to the payload. The bus consists of several subsystems. An artificial satellite (space segment) is presented in Fig. 2 [8].



Fig. 2. Artificial satellite [8]

Every satellite (especially, microsatellite when is dedicated for scientific purposes) carries special instruments that enable it to perform its mission [9] (for example, a satellite that studies the universe has a telescope, a satellite that helps forecast the weather carries cameras to track the movement of clouds). Thus, artificial satellites are classified according to their mission. There are six main types of artificial satellites, classified as follows [10]:

- *Scientific research satellites*
- *Weather satellites*
- *Communications satellites*
- *Navigation satellites*
- *Earth observing satellites*
- *Military satellites*

*Scientific research satellites* gather data for scientific purposes. These satellites during performing their missions gather information about the composition and effects of the space around the Earth, record changes in Earth and its atmosphere and, still others observe planets, stars and other distant objects. Most of these satellites operate in low altitude orbits (LEO). Scientific research satellites also orbit other stars and planets (Mars,

Moon, etc). Usually, these satellites communicate with ground stations in S-band.

*Weather satellites* are dedicated for analyses related to weather forecast. Weather satellites observe the atmospheric conditions over large areas. Their instruments measure cloud cover, temperature, air pressure, precipitation etc. Most of these satellites operate in low altitude orbits (LEO) and in S-band. These satellites always observe the Earth at the same local time. These weather data collected under constant sunlight conditions, then can be easier compared. An example of Europe's image recorded from LEO weather satellite is presented in Fig. 3 [8].

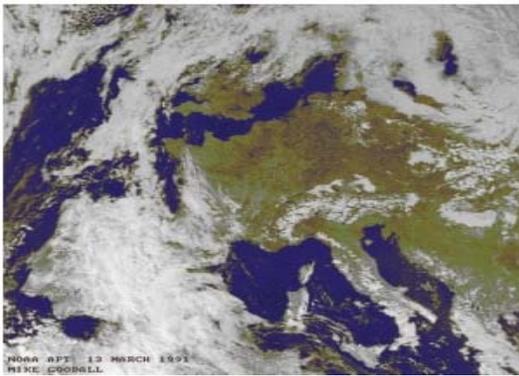


Fig. 3. Image of Europe [8]

*Communication satellites* serve as relay stations, receiving radio signals from one location and transmitting them to another. A communication satellite can relay several television programs or very large number of telephone calls, and data services at once. Communication satellites are usually launched in a high altitude; such is geosynchronous orbit (GEO) [10].

*Navigation satellites* enable operator of aircraft, ships, and land vehicles anywhere on Earth to determine their location with high accuracy. The satellites send out radio signals that are picked up by a computerized receiver carried on an aircraft, ships, or land vehicles. Navigation satellites operate in networks in medium and low Earth orbits (MEO & LEO).

*Earth observing satellites* are used to map and monitor our planet's resources and ever-changing chemical-biological life. They follow LEO orbits. Under constant illumination from the Sun, they take pictures in different colors of visible light and non-visible radiation. Scientists use Earth observation satellites to locate mineral deposits, determine the location and size of freshwater supplies identify sources of pollution and study its effects, etc. [10].

*Military satellites* include weather, communications, navigation and Earth observing satellites used for military purposes.

The scientific research satellites, Earth observing satellites and weather satellites use low Earth orbits (LEO). These satellites accomplish their missions mainly based on photo imagery.

### 3 Orbits

The basic resources available for communication with satellites are *radio frequency spectrum* (RF) and *orbits*. Frequency allocations are treated by international agreements. The orbit is the trajectory followed by the satellite. Several types of orbits are possible, each suitable for a specific application or mission. Generally, the orbits of communication satellites are ellipses within the orbital plane defined by space orbital parameters. Orbits with zero eccentricity are called *circular orbits*. The circularity of the orbit simplifies the analysis. The movement of the satellite within its circular orbit is represented by *orbital time*, *radius*, *altitude* and *velocity*. Circular orbits are categorized based on the altitude above Earth's surface as presented in Fig. 4.

- GEO (Geosynchronous Earth Orbits)
- MEO (Medium Earth Orbits) and
- LEO (Low Earth Orbits)

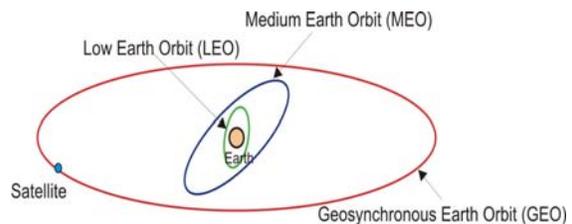


Fig. 4. Satellite orbits

#### 3.1 Low Earth Orbits (LEO)

Low Earth Orbits (LEOs) are just above Earth's atmosphere, where there is almost no air to cause drag on the satellite and reduce its speed. Satellites that point toward deep space dedicated to provide scientific space information generally operate in this type of orbit. The Hubble Space Telescope, for example, operates at an altitude of about 610 km with an orbital period of 97 minutes [10]. LEO altitudes range from 275km up to 1400km limited by Van Allen radiation effects (sensors, integrated circuits and solar cells can be damaged by this radiation) [11]. Satellites in these orbits have an orbital period of around (90-110) minutes. For

satellites this is a short flyover period, which means that the antenna at the ground station must follow the satellite very fast with high pointing accuracy. The contact communication time between the satellite and the ground station takes (5-15) minutes 6-8 times during the day [5], [6]. Mismatch in pointing will lead to a decrease of received signal strength and further to a reduction of the communication quality [4].

### 3.2 Constellation

The constellation is a system of low (medium) Earth orbit (LEO or MEO) identical satellites, launched in several orbital planes with the orbits having the same altitude. The satellites move in a synchronized manner in trajectories relative to Earth. The application of low Earth orbit satellites organized in a *constellation* is an alternative to wireless telephone networks. Satellites in low orbits arranged in a constellation, work together by relaying information to each other and to the users on the ground. If satellites within a constellation are equipped with advanced on-board processing, they can communicate directly with each other by line of sight using inter-satellite links (ISL). If the ISL is between satellites in the same orbit, it is called intra-plane ISL, and if it is between satellites in adjacent planes it is called inter-plane ISL. One of today life constellations is GPS (Global Positioning System) [12]. GPS constellation is presented in Fig. 5 [13].

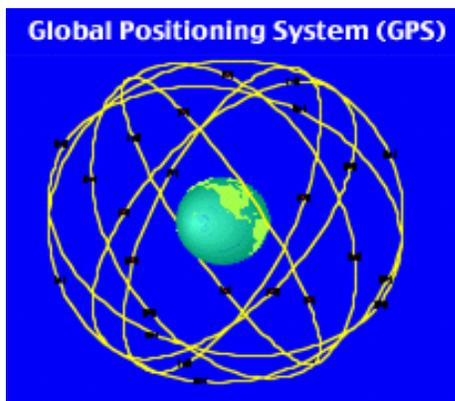


Fig. 5. GPS constellation

This constellation is organized in 6 orbital planes of 4 satellites per plane (24 satellites). Each satellite circles the Earth twice a day [12].

## 4 Space Orbital Parameters

The path of the satellite's motion is an orbit. The orbit is a trajectory within an orbital plane and shaped as an ellipse, with a maximum extension from the Earth center at the *apogee* ( $r_a$ ) and the minimum at the *perigee* ( $r_p$ ) as presented in Fig. 6 [14], [15].

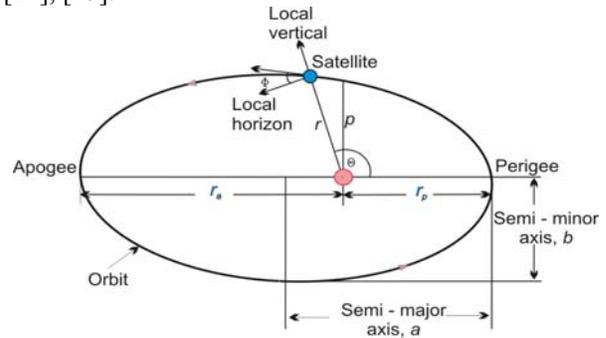


Fig. 6. Major parameters of an elliptical orbit

In order to describe the satellite's movement within its orbit in space, a few parameters are required to be defined. These are known as *space orbital parameters* schematically presented in Fig. 7 and defined under below items a), b), c) and d) [14], [15].

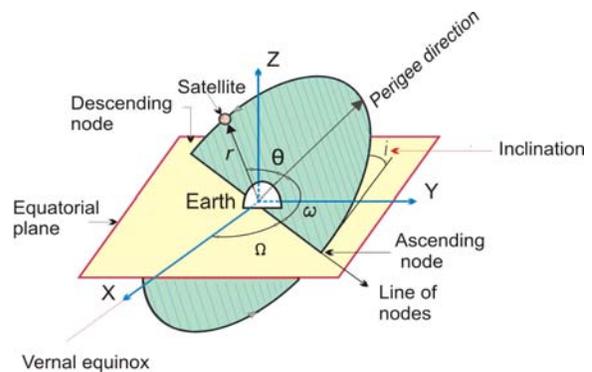


Fig. 7. Space orbital parameters

a) *The position of the orbital plane in space.*

This is specified by means of two parameters - the *inclination*  $i$  and the *right ascension of the ascending node*  $\Omega$ . Inclination  $i$  represents the angle of the orbital plane with respect to the Earth's equator. The right ascension of the ascending node  $\Omega$  defines the location of the ascending and descending orbital crossing nodes (these two nodes make a *line of nodes*) with respect to a fixed direction in space. The fixed direction is Vernal equinox. Vernal equinox is direction of line joining the Earth's center and the Sun on the first day of spring [1], [14], [15].

b) *Location of the orbit in orbital plane.*

Normally an infinite number of orbits can be laid within an orbital plane. So, the orientation of the orbit in its plane is defined by the *argument of perigee*  $\omega$ . This is the angle, taken positively from  $0^\circ$  to  $360^\circ$  in the direction of the satellite's motion, between the direction of the ascending node and the direction of perigee [1], [14], [15].

c) *Position of the satellite in the orbit.*

The position of the satellite in orbit is determined by the angle  $\theta$  called the *true anomaly*, which is the angle measured positively in the direction of satellite's movement from  $0^\circ$  to  $360^\circ$ , between the direction of perigee and the position of the satellite [1], [14], [15].

d) *The shape of orbit.*

The shape of orbit is presented by the *semi-major axis*  $a$  (Fig. 3) which defines the size of orbit and the *eccentricity*  $e$  which defines the shape of the orbit. The eccentricity is the ratio of difference to sum of apogee ( $r_a$ ) and perigee ( $r_p$ ) radii as in Eqn. 1..

$$e = \frac{r_a - r_p}{r_a + r_p} \quad (1)$$

Applying geometrical ellipse features yield out the relations between semi major axis, apogee and perigee as:

$$r_p = a(1 - e) \quad (2)$$

$$r_a = a(1 + e) \quad (3)$$

both,  $r_p$  and  $r_a$  are considered from the Earth's center. Earth's radius is  $r_E = 6378$  km. Then, the highs of perigee and apogee are:

$$h_p = r_p - r_E \quad (4)$$

$$h_a = r_a - r_E \quad (5)$$

For orbits with zero eccentricity, yields:

$$e = 0 \Rightarrow r_a = r_p = a \quad (6)$$

### 5 The Horizon Plane

The Earth rotates from East to West. This is known as eastward direction, the opposite is called westward direction. An orbit in which satellite moves in the same direction as the Earth's rotation is known as *prograde* or *direct orbit*. An orbit in which the satellite moves in opposite direction to Earth rotation is called *retrograde orbit*, as in Fig. 8 is presented [16]. The inclination of a prograde orbit always lies between  $0^\circ$  and  $90^\circ$  (Fig. 7). The inclination of a retrograde orbit always lies between  $90^\circ$  and  $180^\circ$ .

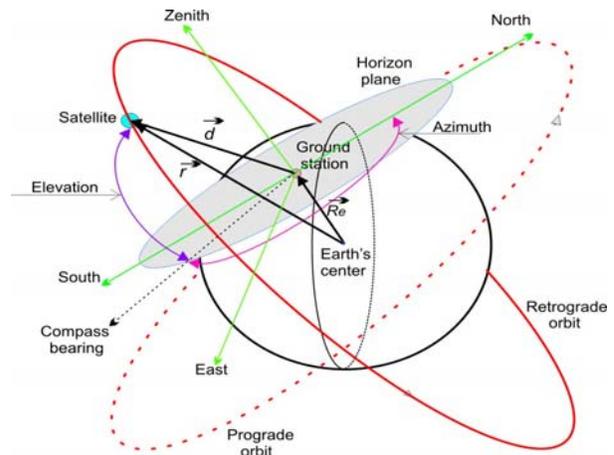


Fig. 8. Orbit and ground station

In Fig. 8, vectors  $\vec{r}$ ,  $\vec{d}$  and  $\vec{R}_e$  represent respectively, vector from the Earth's center to satellite, vector from the ground station to satellite and vector from the Earth's center to ground station. It is obvious the relationship as [15]:

$$\vec{d} = \vec{r} - \vec{R}_e \quad (7)$$

The position of the satellite within its orbit considered from the ground station point of view can be defined by *Azimuth* ( $A_z$ ) and *Elevation* ( $El$ ) angles. The concept of azimuth and elevation taken from Fig 8 is presented in Fig. 9. The azimuth is the angle of the direction of the satellite, measured in the horizon plane from geographical north in clockwise direction. The range of azimuth is  $0^\circ$  to  $360^\circ$ . The elevation is the angle between a satellite and the observer's (ground station's) horizon plane, as presented in Fig. 8, 9. The range of elevation is  $0^\circ$  to  $90^\circ$ .

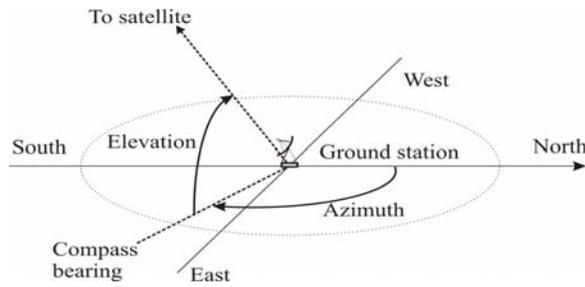


Fig. 9. Azimuth and elevation

The ellipse in Fig. 9 is the ideal horizon plane under elevation of 0° for all azimuths (0-360°).

### 6 Tracking of the Satellite

For tracking the satellite a tracking mechanism and appropriate software is used. As inputs Keplerian elements are used, calculating the actual position of the satellite. The respective software provides real-time tracking information, usually displayed in “radar map” [2], [3]. The display mode of “radar map” includes the accurate satellite’s position with the ground station considered at the center, as in Fig. 10 presented.

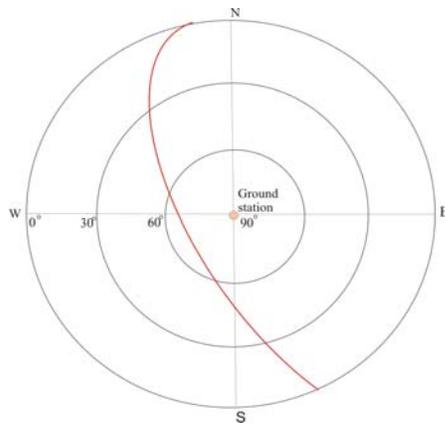


Fig. 10. Radar map presentation

The perimeter of the circle is the horizon plane, with the North on the top ( $Az = 0^\circ$ ), then at the East ( $Az = 90^\circ$ ), South ( $Az = 180^\circ$ ) and West ( $Az = 270^\circ$ ). Three circles represent different elevation  $0^\circ$ ,  $30^\circ$  and  $60^\circ$ . At the center the elevation is  $El=90^\circ$ . Software parameters which define the movement of the satellite related to the ground station are:  $AOS_{time}$  – Acquisition of the satellite (time),  $LOS_{time}$  – Loss of the satellite (time),  $AOS_{Az}$  – Acquisition of the satellite (azimuth),  $LOS_{Az}$  – Loss of the satellite (azimuth),  $Max El$ - Maximal Elevation and  $Orbit$  – Orbit number.

The satellite’s movement (satellite’s pass) is presented with satellite’s path in radar map (red line), what in fact is the satellite’s orbit projection on the horizon plane (Fig. 8). The communication duration is defined as:

$$Duration = AOS_{time} - LOS_{time} \quad (8)$$

and represents the maximum theoretical time duration of the communication between the satellite and ground station [2], [3].

The orbital plane is in principle fixed and defined by orbital parameters (see Fig. 7). Because of Earth’s rotation around its N-S axis for angle  $\beta$ , as depicted in Fig. 11, the ground station changes the position relatively to orbital plane, so the pointing (azimuth and elevation) from the ground station to the satellite is not identical for the both satellite passes (see a) and b) in Fig. 11) [3], [8]. This is represented by picture in Fig. 12, also [17].

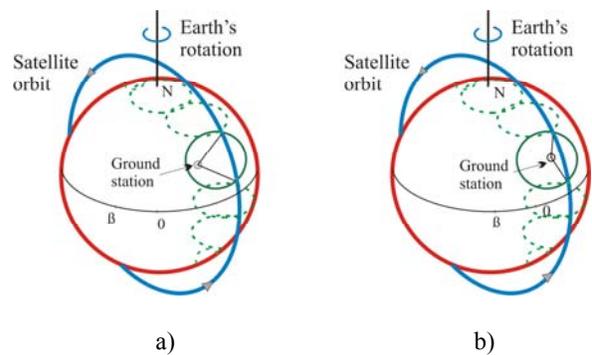


Fig. 11. Satellite pass for an Earth rotation angle of  $\beta$  per orbit a) first pass and b) second pass

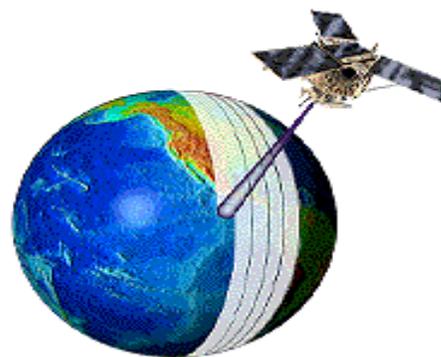


Fig. 12. LEO satellite and coverage area [17]

Because of the Earth’s motion around its North-South axis the satellite passes at the ground station change from pass to pass.

This is typical for LEO satellites which move too fast over the Earth. Because of the Earth's motion the satellite's coverage areas on Earth change, especially for LEO satellites which move too fast over the Earth [18], as it is presented in Fig. 11, 12. Hence the communication duration between the satellite and the ground station is not constant and varies for each orbit path. Each orbit path is characterized with *Max El* [2],[3]. The duration time of communication expressed in Eqn. 8 is based on Kepler's laws and represents the theoretical time duration, considering the acquisition and loss of satellite under  $0^\circ$  elevation [2]. This leads to the fact that there will be many passes with maximum elevation angles below  $5^\circ$ . The communication efficiency due to the time variations as well as the usefulness of low elevation passes will be analyzed for the Vienna satellite ground station. But, practically the acquisition and loss of satellite do not occur at  $0^\circ$  elevation, because of natural barriers or misspointing. So, the communication time depends on one hand on the maximum elevation and on the other hand on the practical radio horizon [2], [3], and [16]. Within this paper the problem of practical horizon and maximal elevation is considered, applying data recorded at the Vienna satellite ground station within the Canadian space observation project "MOST" (Micro variability and Oscillations of Stars) [5], [19].

## 7 MOST Satellite and Vienna Ground Station

Most satellites are launched in a prograde orbit because the Earth's rotational velocity provides part of the orbital velocity with a consequent saving in launch energy [14], [15]. The project "MOST" is a Canadian micro satellite space telescope mission. The size of the satellite is 65cm x 65cm x 30cm and the mass is about 65kg. The goals of the mission are: to analyze the inner structure of stars, set a lower limit to the age of the universe and to search for Exoplanets, by picking up tiny light variations of stars. The project "MOST" consists of a Low Earth Orbiting (LEO) Satellite and three Ground Stations, one of them in Vienna [5], [19]. The idea of "MOST" satellite is depicted in Fig. 13.

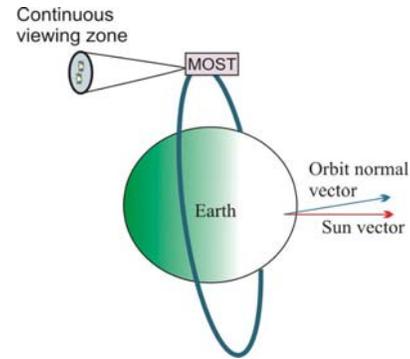


Fig. 13. MOST satellite idea

The baseline orbit of MOST is a sun-synchronous orbit, with  $98^\circ$  inclination and an altitude of around 820 km. The Vienna ground station system was set up at the Institute for Astronomy of the University of Vienna in cooperation with the Institute of Communications and Radio-Frequency Engineering of the Vienna University of Technology. The ground station must track the satellite during its flyover keeping a pointing accuracy of  $0.5^\circ$  [5], [19]. The ground station can interact with the satellite only if it is visible above the horizon and therefore for a fraction of few orbits per day [19]. The visibility region of the Vienna satellite ground station is shown in Fig. 14.

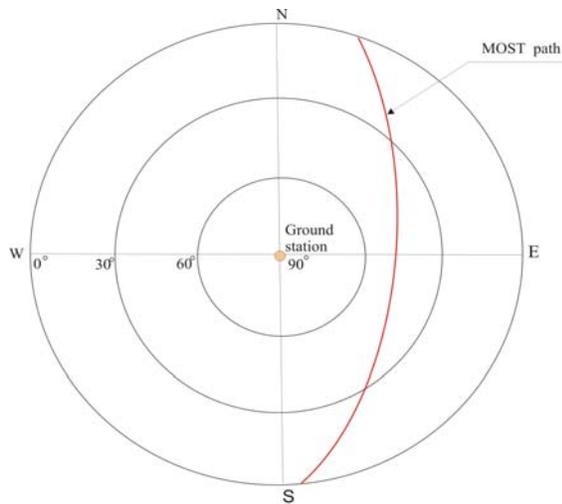


Fig. 14. Visibility region of Vienna ground station for elevation angle of  $0^\circ$  [20]

The MOST satellite has a line of sight radio contact with the Vienna ground station 6-8 times per day. The communication duration time of each satellite pass will last between 5-15 minutes [2], [3], [19].

### 7.1 Tracking of the MOST Satellite

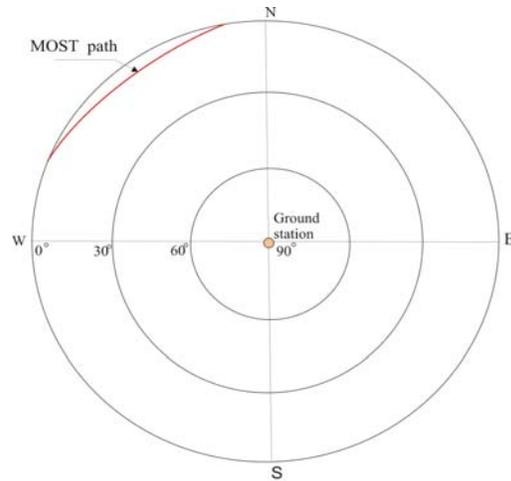
For tracking the satellite a tracking mechanism and software is used. The tracking software used for “MOST” is called *Nova* [20]. As inputs Keplerian elements are used calculating the actual position of the satellite. The respective software provides real-time tracking information. The display mode “radar map” includes the stars with accurate position on the sky with the ground station at the center. In Fig. 15 and Fig. 16 two different MOST satellite passes are presented. These passes are recorded in Vienna and presents the satellite passes in the radar map. The satellite movement in these figures is indicated as “MOST path” [3]. During the first pass shown in Fig. 15 contact between the satellite and ground station was established.



AOS time	05h 48m 58s
LOS time	06h 04m 06s
Duration	00h 15m 06s
AOS Az.	19°
Max. El.	45°
LOS Az.	178°
Orbit	2860

Fig.15. MOST path (1)

At the second pass shown in Fig. 16 no contact to the satellite could be established because of natural barriers at this low elevation. So, the communication time depends on one hand on the maximum elevation and on the other hand on the practical radio horizon. Let us firstly consider the problem of practical horizon plane.

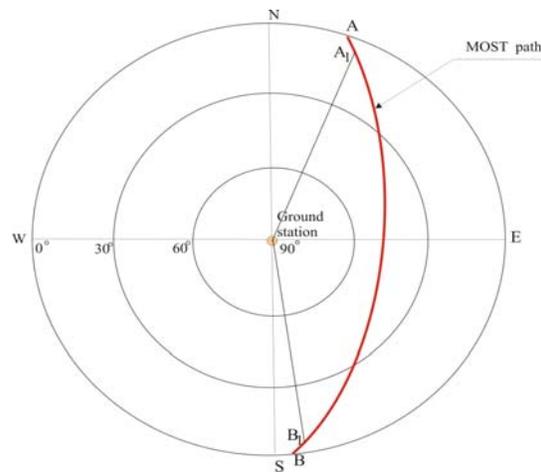


AOS time	09h 10m 47s
LOS time	09h 19m 32s
Duration	00h 08m 44s
AOS Az.	354°
Max. El.	3°
LOS Az.	283°
Orbit	2862

Fig. 16. MOST path (2).

### 7.2 Practical Horizon Plane

In Fig.17, one of the MOST satellite paths with respective data is presented [16].



AOS time	05h 48m 58s
LOS time	06h 04m 06s
Duration	00h 15m 06s
AOS Az.	19°
Max. El.	45°
LOS Az.	178°
Orbit	2860

Fig. 17. MOST path

Theoretically, based on Kepler’s laws, the communication between the satellite and the ground station should be established at point  $A$  ( $Az = 19^\circ$  under  $0^\circ$  elevation) and communication should be lost at point  $B$  ( $Az = 178^\circ$  under  $0^\circ$  elevation) at Fig. 17. No contact to the satellite could be established under the  $0^\circ$  elevation, because of natural barriers or misspointing. Practically contact between the satellite and ground station is established at point  $A_1$  and lost at point  $B_1$  (both under elevation of few degrees). Thus, points  $A$  and  $B$  belong to ideal horizon plan, otherwise  $A_1$  and  $B_1$  belong to practical horizon. In order to create better picture of practical horizon at Vienna satellite ground station appropriate data about acquisition ( $AEI$ ) and lost elevation ( $LEI$ ) for different orbits are recorded. Few of these records are presented in Table 1 [16].

Table 1. Orbit records at Vienna ground station

Orbit	AAz	AEI	MaxEl	LAz	LEI
xxxx	°	°	°	°	°
2639	170	1.0	72	340	2.5
2652	107	1.0	16	7	5.0
2662	350	4.0	10	270	5.0
2676	352	4.0	14	258	3.0
2709	122	1.0	22	2	5.0
2737	95	1.0	12	8	5.0
2759	49	6.0	13	139	1.0
2760	15	4.0	84	190	4.0
2817	36	6.0	21	154	1.0
2879	103	1.5	14	14	7.0
2907	61	6.0	7	20	4.5
2945	26	5.0	34	163	7.0
2947	347	3.0	7	276	1.5
2965	149	2.0	56	350	2.0
2966	206	4.0	24	331	3.0
2987	43	4.0	11	129	3.0
3009	236	3.5	10	315	4.0
3066	253	2.0	5	302	4.0
3072	79	7	8	118	3.0
3136	187	3.5	41	337	3.5
3160	339	3.0	6	290	2.5
3193	203	5.5	27	332	3.0
3251	215	3.0	17	327	3.0
3302	324	4.5	5	300	3.0
2992	65	2.0	5	23	4.0

Data recorded at Vienna ground station (included those from Table 1) are pointed in radar map view, presented in Fig. 18 [16].

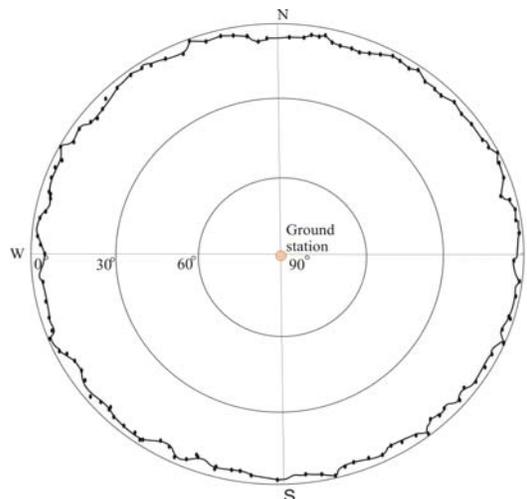


Fig. 18. Practical horizon plane

The inner line (darker one) related to the largest circle is in fact the *practical horizon plane*. It is obvious that the practical horizon is always shorter than ideal one, reflecting on shorter communication time between the satellite and the ground station.

### 7.3 Communication Duration

The duration time of communication expressed in Eqn. 8 is based on Kepler’s law and represents the theoretical time duration, which practically is always shorter because of natural barriers.

Table 2. Expected orbital data

Orbit	AOStime	LOStime	Duration	MaxEl
xxxx	hh:mm:ss	hh:mm:ss	mm:ss	°
2235	5:31:39	5:47:10	15: 21	84.0
2378	7:21:05	7:33:23	12: 18	14.0
2467	13:36:00	13:45:38	9: 38	6.0
2795	16:11:18	16:26:47	15: 29	86.0
2893	14:00:22	14:11:27	11: 05	11.0
2965	15:41:33	15:56:47	15: 14	56.0
3007	14:45:14	14:58:50	13: 36	22.0
3009	18:07:10	18:18:41	11: 31	9.0
3093	16:09:01	16:24:30	15: 29	90.0
3160	8:08:24	8:17:27	9: 03	8.0
3263	14:02:10	14:13:25	9: 15	10.0
3302	8:17:03	8:25:20	8: 17	5.0
3336	18:24:34	18:38:41	14: 07	23.0
3427	3:43:08	3:55:43	12 : 35	13.0

Table 2, shows examples for the *AOS<sub>time</sub>*, *LOStime* and *MaxEl* taken from MOST passes at the Vienna ground station. The different time duration for each orbit in Table 2, based on Eqn. 8 confirms the explanation by the Fig. 11 and Fig. 12. For practical reasons and in order to have real view on communication time, the *LOCKtime* (time when communication was established) and the *UNLOCKtime* (time when communication was lost) with its respective Acquisition Elevation (*AEI*) and Lost Elevation (*LEI*) was recorded and is presented in Table 3 [3].

Table 3. Real orbital data

Orbit	LOCK time	UNLOCK time	AEI	LEI
xxxx	hh:mm:s	hh:mm:s	°	°
2235	5:32:32	5:47:00	3.0	2.0
2378	7:22:45	7:32:43	4.0	2.5
2467	13:38:20	13:43:20	4.0	4.0
2795	16:11:20	16:26:00	2.0	2.0
2893	14:01:00	14:10:00	2.0	3.5
2965	15:42:00	15:56:00	2.0	2.0
3007	14:46:00	14:57:00	3.0	4.5
3009	18:08:00	18:17:00	3.5	4.0
3093	16:10:00	16:25:00	2.0	2.5
3160	8:10:00	8:16:00	3.5	2.5
3263	14:03:00	14:12:00	3.0	3.5
3302	8:20:00	8:23:00	4.0	3.0
3336	18:25:00	18:37:00	3.0	2.5
3427	3:44:00	3:54:00	3.0	3.0

Usually the lock is established and lost in average at elevation angles of 1° – 4°. In order to quantify these variations in communication duration (comparing ideal and real communication time), a parameter called *Time Efficiency Factor (T<sub>ef</sub>)* was defined [2], [3]:

$$T_{ef} = \frac{LOCK_{time} - UNLOCK_{time}}{AOS_{time} - LOS_{time}} \quad (9)$$

The *Time Efficiency Factor* represents the ratio of the real communication duration to the theoretical communication duration. Fig. 19 shows *T<sub>ef</sub>* in percent as function of *MaxEl* by using the data from Table 2 and Table 3. From the Fig. 19 it is obvious, that for *MaxEl* higher than 10° it is the time variance which is keeping a trend of linearity starting at 80% toward 100%, but for *MaxEl* lower than 10° *T<sub>ef</sub>* rapidly falls causing high time variance.

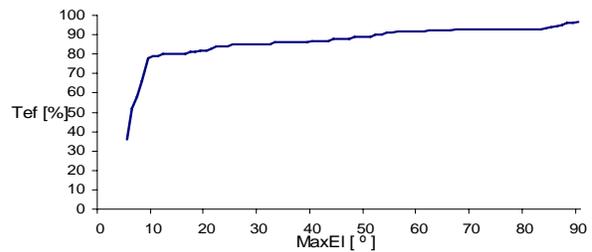


Fig. 19. Time efficiency factor dependency on maximal elevation angle

Assuming a pass with an elevation of 5° with practical time duration of 3 min. and a data rate of 38.4 kbit/s, the amount of data (including the protocol) which can be downloaded during this pass is 863 kByte. Further assuming a protocol overhead of about 15% the data amount downloaded during this pass is about 735 kByte. This is worth amount of data collected by the satellite during low elevation. This confirms the necessity of careful design under low elevation, also.

## 8 Conclusions

Motion detection enables practical horizon determination. Obviously practical horizon differs from the ideal one, for at least 2°-3° degrees of elevation in average, because of natural barriers. This is confirmed based on records at Vienna satellite ground station. Considering that for LEO satellites the contact between satellite and ground station is in range of (3-15) minutes, this difference impacts communication duration, also. During communication with LEO satellites it is obvious that at elevation angles below 10° the time efficiency falls, because of natural barriers, thus not the whole pass can be used. This leads to a decreased data flow compared to the theoretical case. The analysis of the data amount at a low elevation pass has shown that it is worth to dimension the ground station also for low elevation passes, because an important part of the stored data at the satellite can be downloaded at such passes. Finally, Time Efficiency Factor could be considered as a QoS element on communication duration between a satellite and a ground station.

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## References:

- [1] G. Maral, M. Bousquet, *Satellite Communications Systems*, John Wiley & sons, UK, 2005.
- [2] S. Cakaj, K. Malaric, Rigorous analysis on performance of LEO satellite ground station in urban environment, *International Journal Satellite Communications and Networking*, Vol.25, No.6, 2007, pp. 619-643.
- [3] S. Cakaj, W. Keim, K. Malaric, Communication duration with low Earth orbiting satellites, *Proc. IEEE, IASTED, 4<sup>th</sup> International Conference on Antennas, Radar and Wave Propagation*, Montreal, Canada, 2007, pp. 85-88.
- [4] E. Alejandro et al., Control system for the antenna positioning, *Proc. WSEAS, 7<sup>th</sup> International Conference on Applied Informatics and Communications*, Athens, Greece, 2007, pp. 390 - 392.
- [5] R.E. Zee, P. Stibrany, The MOST Microsatellite: A low – cost enabling technology for future space science and technology missions, *Canadian Aeronautics and Space Journal*, Vol. 48, No. 1, Canada, 2002, pp. 1-10.
- [6] W. Keim, V. Kudielka, A.L. Scholtz, A Satellite ground station for an urban environment, *Proc., IASTED, International Conference on Communication Systems and Networks*, Marbella, Spain, 2004, pp. 280-284.
- [7] M. Reyes- Ayala et al., UW Detection in TDMA satellite system, *Proc. WSEAS, 6<sup>th</sup> International Conference on Applied Informatics and Communications*, Elounda, Greece, 2006, pp. 237 - 240.
- [8] Radio Frequency Spectrum Management Course, USTTI (United States Telecommunication Training Institute) course M6 -102, Washington DC, 2006.  
<http://www.noaa.gov>
- [9] E.A. Essex et al., Monitoring the ionosphere/ plasmasphere with Low Earth Orbit satellites: The Australian microsatellite FedSat, *Cooperative Research Center for Satellite Systems*, Department of Physics, La Trobe University, Bundoora, Australia.
- [10] J.E. Oberright, *Satellite artificial*, World Book Online Reference Center, World Book, Inc 2004.
- [11] <http://www.answers.com/topic/van-allen-radiation> - belt
- [12] N. Enescu, D. Mancas, E. Manole, Locating GPS coordinates on PDA, *Proc. WSEAS, 8<sup>th</sup> International Conference on Applied Informatics and Communications*, Rhodes, Greece, 2008, pp. 470 - 474.
- [13] <http://www.gps.gov>
- [14] M. Richharia, *Satellite Communications Systems*, McGraw Hill, New York, 1999.
- [15] D. Roddy, *Satellite communications*, McGraw Hill, New York, 2006.
- [16] S.Cakaj, M. Fischer, A.L. Scholtz, Practical horizon plane for low Earth Orbiting (LEO) satellite ground stations, *Proc. WSEAS, 9<sup>th</sup> International Conference on Applied Informatics and Communications*, Istanbul, Turkey, 2009, pp. xx – xx.
- [17] Radio Frequency Spectrum Management Course, USTTI (United States Telecommunication Training Institute) course M6 -102, Washington DC, 2006.  
<http://www.thetech.org/exhibits/online/satellite>
- [18] B. Elbert, *Ground Segment and Earth Station Handbook*, Artech House Publishers, Boston, 2001.
- [19] W. Weis, W. Keim, The Austrian contribution to: Microvariability and Oscillation of Stars – Detection of Exoplanets, *Report, Universitet Wien, Department of Astronomy*, Vienna, Austria 2003.
- [20] Northern Lights Software Associates,  
<http://www.nlsa.com>