# Analytic Solution of Nadir Attitude Pointing for LEO Microsatellite 

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#### Abstract

An Analytic solution of Nadir attitude pointing equation of gravity gradient satellite stabilised is presented. The attitude equation is Euler linearised equation for near Nadir pointing axially symmetric satellite including only gravity gradient torque and assuming other torques such as magnetic torque, aerodynamic torque, solar radiation pressure torque and controller are constants. The obtained analytical solution was compared to numerical solution of satellite attitude equation.


Keywords:- LEO, Microsatellite, Alsat-1, Analytic, Nadir, Attitude, Simulation.

## 1 Introduction

A spacecraft in orbit always needs to stabilise the attitude against the external disturbance torques acting on it. Attitude control usually needs to be autonomous or semi-autonomous. On Alsat-1 [1], [6] the available actuators are reaction/momentum wheels and magnetic torquing. A mixture of attitude estimation and control algorithms is needed: these take the sensor measurements as inputs, compute the attitude and rates of the satellite, and then send commands to the actuators to maintain or stabilise that attitude, or direct the satellite to a new attitude.
Alsat-1 exploited the passive gravity gradient torque [4]. A substantial amount of literature has studied the technical problems of ADCS in many different areas.
The motion of a spacecraft presents two dynamic aspects of interest. Classical dynamics allows, under certain general conditions, for the motion of a body to be treated as the combination of two motions: a translational motion of the centre of mass and a rotation of the body about the centre of mass. The theory of attitude control generally considers only
the second effect and ignores the first. The application of any force can only be interpreted as the resultant torque that would exist around the centre of mass and ignores any change to the translational velocity [8].
The equations of motion of a spacecraft can be divided into two parts: The dynamic equations of motion and kinematic equations of motion. The dynamic equations of motion express the relationship between the spacecraft body angular rate and the applied torque. These are necessary for dynamic simulations and for attitude prediction, whenever gyroscopic measurements of the angular rate is unavailable. The kinematic equations of motion are a set of first-order differential equations expressing the relationship between the attitude parameters and the rate [9].

## 2 Analytical Solution:

From [2], [3] and [5] the linearised Euler equation for near Nadir pointing axially symmetric satellite is given as follows

$$
\begin{align*}
& \ddot{\phi}+4(1-\mathrm{k}) \omega_{0}^{2} \phi-\mathrm{k} \omega_{0} \dot{\psi}=\mathrm{n}_{\mathrm{x}}  \tag{1.a}\\
& \ddot{\theta}+3(1-\mathrm{k}) \omega_{0}^{2} \theta=\mathrm{n}_{\mathrm{y}}  \tag{1.b}\\
& \ddot{\psi}+\omega_{0} \dot{\phi}=\mathrm{n}_{\mathrm{z}} \tag{1.c}
\end{align*}
$$

Initial conditions is given as follows

$$
\begin{align*}
& \phi\left(\mathrm{t}_{0}\right)=\phi_{0}, \theta\left(\mathrm{t}_{0}\right)=\theta_{0}, \psi\left(\mathrm{t}_{0}\right)=\psi_{0} \\
& \dot{\phi}\left(\mathrm{t}_{0}\right)=\dot{\phi}_{0}, \dot{\theta}\left(\mathrm{t}_{0}\right)=\dot{\theta}_{0}, \dot{\psi}\left(\mathrm{t}_{0}\right)=\dot{\psi}_{0} \tag{2}
\end{align*}
$$

Where
$\omega_{0} \quad$ : orbital rate;
$\theta \quad$ : roll angle in rad ;
$\varphi \quad$ : pitch angle in rad ;
$\psi \quad$ : yaw angle in rad;
$\mathbf{I}=\operatorname{diag}\left[I_{T} I_{T} I_{z}\right]$ moment of inertia tensor of the

$$
\text { Spacecraft, } \mathrm{k}=\frac{\mathrm{I}_{\mathrm{z}}}{\mathrm{I}_{\mathrm{T}}}
$$

$\mathbf{N}=\left[\mathrm{n}_{\mathrm{x}} \mathrm{n}_{\mathrm{y}} \mathrm{n}_{\mathrm{z}}\right]^{\mathrm{T}}$ normalized torque induced by controller or unmodelled disturbances torque.

We want to find tout he analytic solution of Equation (1) by assuming $n_{x}, n_{y}$ and $n_{z}$ constants. In order to make the solution simpler $\mathrm{k} \ll 1$ is assumed which is true for Alsat-1 microsatellite [3].

### 2.1 Pitch Equation:

The solution of the pitch equation is given as follows
$\theta(\mathrm{t})=\frac{\mathrm{n}_{\mathrm{y}}}{3 \omega_{0}^{2}}(1+\mathrm{k})+\mathrm{A}_{1} \cos \theta \omega_{\theta} \mathrm{t}+\mathrm{A}_{2} \sin \theta \omega_{\theta} \mathrm{t}$
Where
$\omega_{\theta}^{2}=3(1-\mathrm{k}) \omega_{0}^{2}$
$\mathrm{A}_{1}, \mathrm{~A}_{2}$ integral constant.
From the initial condition, the solution of the pitch equation will be

$$
\begin{align*}
\theta(\mathrm{t})= & \frac{\mathrm{n}_{\mathrm{y}}}{3 \omega_{0}^{2}}(1+\mathrm{k})+\left(\theta_{0}-(1+\mathrm{k}) \frac{\mathrm{n}_{\mathrm{y}}}{3 \omega_{0}^{2}}\right) \cos \theta \omega_{\theta} \mathrm{t}  \tag{4}\\
& +\left(1+\frac{\mathrm{k}}{2}\right) \frac{\dot{\theta}_{0}}{\sqrt{3} \omega_{0}} \sin \theta \omega_{\theta} \mathrm{t}
\end{align*}
$$

### 2.2 Roll Equation:

Equation (1.c) can be integrated as

$$
\begin{equation*}
\dot{\psi}+\omega_{0} \phi=\mathrm{n}_{\mathrm{z}} \mathrm{t}+\mathrm{A}_{3} \tag{5}
\end{equation*}
$$

From the initial condition, the above equation will take the following form

$$
\begin{equation*}
\dot{\psi}+\omega_{0} \phi=\mathrm{n}_{\mathrm{z}} \mathrm{t}+\dot{\psi}_{0}+\omega_{0} \phi_{0} \tag{6}
\end{equation*}
$$

Substituting equation (6) in roll equation (1.b) yields

$$
\begin{equation*}
\ddot{\phi}+(4-3 \mathrm{k}) \omega_{0}^{2} \phi=\mathrm{n}_{\mathrm{x}}+\mathrm{k} \omega_{0}^{2}\left(\frac{\dot{\psi}_{0}}{\omega_{0}}+\phi_{0}\right)+\mathrm{k} \omega_{0} \mathrm{n}_{\mathrm{z}} \mathrm{t} \tag{7}
\end{equation*}
$$

Regarding to the initial condition, the solution is as follows

$$
\begin{align*}
\phi & =\frac{1}{4}\left(1+\frac{3}{4} \mathrm{k}\right) \frac{\mathrm{n}_{\mathrm{x}}}{\omega_{0}^{2}}+\frac{\mathrm{k}}{4}\left(\frac{\dot{\psi}_{0}}{\omega_{0}}+\phi_{0}\right)+\frac{\mathrm{k}}{4} \frac{\mathrm{n}_{\mathrm{z}}}{\omega_{0}} \mathrm{t}+ \\
& {\left[\phi_{0}-\frac{1}{4}\left(1+\frac{3}{4} \mathrm{k}\right) \frac{\mathrm{n}_{\mathrm{x}}}{\omega_{0}^{2}}-\frac{\mathrm{k}}{4}\left(\frac{\dot{\psi}_{0}}{\omega_{0}}+\phi_{0}\right)\right] \cos \omega_{\phi}^{\mathrm{t}}+}  \tag{8}\\
& {\left[\frac{1}{2}\left(1+\frac{3}{8} \mathrm{k}\right) \frac{\dot{\phi}_{0}}{\omega_{0}}-\frac{\mathrm{k}}{8} \frac{\mathrm{n}_{\mathrm{z}}^{2}}{\omega_{0}^{2}}\right] \sin \omega_{\phi}^{\mathrm{t}} }
\end{align*}
$$

### 2.3 Yaw Equation:

Substituting roll equation (8) into equation (5), yields

$$
\begin{aligned}
& \psi=\psi_{0}-\frac{1}{4}\left(\left(1+\frac{3}{4} \mathrm{k}\right) \frac{\dot{\phi}_{0}}{\omega_{0}}-\frac{\mathrm{k}}{4} \frac{\mathrm{n}_{\mathrm{z}}^{2}}{\omega_{0}^{2}}\right) \\
& -\left[\frac{1}{4}\left(1+\frac{3}{4} \mathrm{k}\right) \frac{\mathrm{n}_{\mathrm{x}}}{\omega_{0}^{2}}-\left(1-\frac{\mathrm{k}}{4}\right)\left(\frac{\dot{\psi_{0}}}{\omega_{0}}+\phi_{0}\right)\right] \omega_{0} \mathrm{t} \\
& -\frac{1}{2}\left[\left(1+\frac{3}{8} \mathrm{k}\right) \phi_{0}-\frac{1}{4}\left(1+\frac{9}{8} \mathrm{k}\right) \frac{\mathrm{n}_{\mathrm{x}}}{\omega_{0}^{2}}-\frac{\mathrm{k}}{4}\left(\frac{\dot{\psi}_{0}}{\omega_{0}}+\phi_{0}\right)\right] \sin \omega_{\phi} \mathrm{t} \\
& +\frac{1}{4}\left[\left(1+\frac{3}{4} \mathrm{k}\right) \frac{\dot{\phi}_{0}}{\omega_{0}}-\frac{\mathrm{k}}{4} \frac{\mathrm{n}_{\mathrm{z}}}{\omega_{0}^{2}}\right] \cos \omega_{\phi} \mathrm{t}+\frac{1}{2}\left(1-\frac{\mathrm{k}}{4}\right) \mathrm{n}_{\mathrm{z}} \mathrm{t}^{2}
\end{aligned}
$$

## 3 Numerical Solution:

From [5], [7] and [10] the satellite attitude equation is given as follows

$$
\left\{\begin{array}{l}
\dot{\omega}_{\mathrm{x}}=\frac{\mathrm{N}_{\mathrm{x}}^{\mathrm{MT}}}{\mathrm{I}_{\mathrm{x}}}+\alpha\left(3 \omega_{0}^{2} \mathrm{a}_{23} \mathrm{a}_{33}-\omega_{\mathrm{y}} \omega_{\mathrm{z}}\right) \\
\dot{\omega}_{\mathrm{y}}=\frac{\mathrm{N}_{\mathrm{y}}^{\mathrm{MT}}}{\mathrm{I}_{\mathrm{y}}}+\beta\left(3 \omega_{0}^{2} \mathrm{a}_{13} \mathrm{a}_{33}-\omega_{\mathrm{x}} \omega_{\mathrm{z}}\right) \\
\dot{\omega}_{\mathrm{z}}=\frac{\mathrm{N}_{\mathrm{z}}^{\mathrm{MT}}}{\mathrm{I}_{\mathrm{z}}}+\gamma\left(3 \omega_{0}^{2} \mathrm{a}_{13} \mathrm{a}_{23}-\omega_{\mathrm{x}} \omega_{\mathrm{y}}\right) \\
\left\{\begin{array}{l}
\dot{\mathrm{q}}_{1} \\
\dot{\mathrm{q}}_{2} \\
\dot{\mathrm{q}}_{3} \\
\dot{\mathrm{q}}_{4}
\end{array}\right]=\frac{1}{2}\left[\begin{array}{ccc}
0 & \omega_{\mathrm{oz}} & -\omega_{\mathrm{oy}} \\
-\omega_{\mathrm{oz}} & \omega_{\mathrm{ox}} \\
\omega_{\mathrm{oy}} & \omega_{\mathrm{ox}} & \omega_{\mathrm{oy}} \\
-\omega_{\mathrm{ox}} & -\omega_{\mathrm{ox}} & 0 \\
\omega_{\mathrm{oz}} \\
-\omega_{\mathrm{oz}} & 0
\end{array}\right]\left[\begin{array}{l}
\mathrm{q}_{1} \\
\mathrm{q}_{2} \\
\mathrm{q}_{3} \\
\mathrm{q}_{4}
\end{array}\right] \\
\omega_{\mathrm{LO}}=\omega_{\mathrm{B}}-\mathbf{A} \omega_{0}  \tag{10}\\
\theta=\arctan \left\{\begin{array}{l}
\left.\frac{\mathrm{a}_{31}}{\mathrm{a}_{33}}\right\} ; \phi=\arcsin \left\{-\mathrm{a}_{23}\right\}
\end{array}\right\} \\
\psi=\arctan \left\{\frac{\mathrm{a}_{12}}{\mathrm{a}_{22}}\right\}
\end{array}\right\} \begin{aligned}
& \dot{\phi}=\omega_{\mathrm{x}} \cos \psi-\omega_{\mathrm{y}} \sin \psi \\
& \dot{\theta}=\left(\omega_{\mathrm{x}} \sin \psi+\omega_{\mathrm{y}} \cos \psi\right) \sec \phi \\
& \dot{\psi}=\omega_{\mathrm{z}}+\left(\omega_{\mathrm{x}} \sin \psi+\omega_{\mathrm{y}} \cos \psi\right) \tan \phi
\end{aligned}
$$

Where
$\alpha=\frac{\mathrm{I}_{\mathrm{z}}-\mathrm{I}_{\mathrm{y}}}{\mathrm{I}_{\mathrm{x}}}, \beta=\frac{\mathrm{I}_{\mathrm{x}}-\mathrm{I}_{\mathrm{z}}}{\mathrm{I}_{\mathrm{y}}}, \gamma=\frac{\mathrm{I}_{\mathrm{y}}-\mathrm{I}_{\mathrm{x}}}{\mathrm{I}_{\mathrm{z}}}$
$\mathbf{N}=\left[\mathrm{N}_{\mathrm{x}}^{\mathrm{MT}} \mathrm{N}_{\mathrm{y}}^{\mathrm{MT}} \mathrm{N}_{\mathrm{z}}^{\mathrm{MT}}\right]^{\mathrm{t}}$ applied magnetorquer control firing;
$\omega_{\mathrm{B}}=\left[\omega_{\mathrm{x}} \omega_{\mathrm{y}} \omega_{\mathrm{z}}\right]^{\mathrm{t}}$ Euler body rates wrt. inertial;
$\omega_{\mathrm{LO}}=\left[\begin{array}{lll}\omega_{\mathrm{ox}} & \omega_{\mathrm{oy}} & \omega_{\mathrm{oz}}\end{array}\right]^{\mathrm{t}}$ Euler local orbital rates wrt. inertial;
$\mathrm{a}_{\mathrm{ij}}$ is an $(\mathrm{i}, \mathrm{j})$ component of DCM matrix $\mathbf{A}$.

## 4 Simulations Results:

The following initialization parameters were utilized

## Normalized Torque

$\mathrm{n}_{\mathrm{x}}=\mathrm{n}_{\mathrm{y}}=\mathrm{n}_{\mathrm{z}}=0$

## Inertial Tensor (Satellite configuration I)

I $\left[\mathrm{kgm}^{2}\right]$
: diag [185 158 5] ${ }^{\mathrm{t}}$

## Miscellaneous

Simulation time [orbit] :2
Integration step [sec] :1
Sampling time $[\mathrm{sec}] \quad: 5$


Fig. 1 Roll Angle


Fig. 2 Pitch Angle


Fig. 3 Yaw Angle


Fig. 4 Roll Rate Angle


Fig. 5 Pitch Rate Angle


Fig. 6 Yaw Rate Angle

Table 1: Lists the angular error

|  | Roll <br> [degree] | Pitch <br> [degree] | Yaw <br> [degree] |
| :--- | :---: | :---: | :---: |
| Average | -0.108 | -0.039 | 0.353 |
| STD $[1-\sigma]$ | 0.142 | -0.395 | -0.362 |
| RMS | 0.178 | 0.397 | 0.506 |

Table 2: Lists the rate error

|  | Roll Rate <br> $[$ deg/sec] | Pitch Rate <br> $[$ deg/sec] | Yaw Rate <br> $[\mathrm{deg} / \mathbf{s e c}]$ |
| :--- | :---: | :---: | :---: |
| Average | $2.9^{*} 10^{-4}$ | $-9.84^{*} 10^{-6}$ | $48^{*} 10^{-4}$ |
| STD $[1-\sigma]$ | $-5.68^{*} 10^{-5}$ | $-7.36^{*} 10^{-4}$ | $1.68^{*} 10^{-4}$ |
| RMS | $2.96^{*} 10^{-4}$ | $7.36^{*} 10^{-4}$ | $49^{*} 10^{-4}$ |

Table 3: Lists the error magnitude angles and rates

|  | Mag Error <br> Average | Mag Error <br> STD $[1-\sigma]$ | Mag Error <br> RMS |
| :--- | :---: | :---: | :---: |
| Angles <br> [deg] | 0.371 | 0.554 | 0.667 |
| Rate <br> $[$ deg/sec] $]$ | $48^{*} 10^{-4}$ | $7.5^{*} 10^{-4}$ | $49.6^{*} 10^{-4}$ |

For the graphs presented above, notice that the magnitude of the RMS error results indicates that the angular error is approximately 0.66 degree and the rate error is about 0.005 degree/second.

## 5 Conclusion

This paper detailed the analytic solution of Nadir attitude pointing equation of gravity gradient LEO satellite.
The magnitude of the RMS error results indicates that the angular error is approximately 0.66 degree and the rate error is about 0.005 degree/second, both in degrees.
A low cost method of full satellite attitude propagator was proposed to be used for LEO microsatellite gravity gradient stabilised (small libration).
The version presented is only valid for LEO microsatellite gravity gradient stabilised including gravity gradient disturbance. The extension to aerodynamic disturbances is in progress.

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